

Which electronic structure method? An introduction to Density Functional Theory (part 1)

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Elementary Quantum Mechanics

- The Schrodinger Equation
- Variational Principle
- The Hartree Fock Approximation
- Correlation energy

Density Functional Theory

- Thomas-Fermi model
- The Hohenberg-Kohn theorems
- Orbitals and Kohn-Sham equations

Recommended readings:

- ▶ *Density-Functional Theory of Atoms and Molecules* by Robert G. Parr, Yang Weitao - Oxford Science Publications
- ▶ *Modern Quantum Chemistry. Introduction to Advanced Electronic Structure Theory* by Attila Szabo and Neil S. Ostlund - Dover Publications
- ▶ *Ab initio Molecular dynamics. Basic Theory and Advanced Methods* by Dominik Marx and Jürg Hutter - Cambridge University Press

The Schrodinger Equation

$$\hat{H}\Psi = E\Psi \quad (1)$$

$$\Psi = \Psi(x_1, x_2, \dots, x_n) \quad (2)$$

$$\hat{H} = \hat{T} + \hat{V}_{ne} + \hat{V}_{ee} \quad (3)$$

$$\hat{H} = \sum_{i=1}^N \left(-\frac{1}{2} \nabla_i^2\right) + \sum_{i=1}^N v(r_i) + \sum_{i<j}^N \frac{1}{r_{ij}} \quad (4)$$

where

$$v(r_i) = - \sum_{\alpha} \frac{Z_{\alpha}}{r_{i\alpha}} \quad (5)$$

- ▶ The total energy is $W = E + V_{nn}$
 where $V_{nn} = \sum_{\alpha<\beta} \frac{Z_{\alpha}Z_{\beta}}{R_{\alpha\beta}}$

The Schrodinger Equation

$$\hat{H}\Psi = E\Psi$$

must be solved subject to appropriate boundary conditions, namely $\Psi(x_1, \dots, x_n)$ decays to zero at infinity, or appropriate periodic boundary condition for a solid.

$$|\Psi(x_1, \dots, x_n)|^2 dr_1 \dots dr_N \quad (6)$$

is the probability of finding the system with position coordinates between x_1, \dots, x_N and $x_1 + dx_1 \dots x_N + dx_N$.

Ψ must be antisymmetric.

$$\int \Psi_k^* \Psi_l dx^N = \langle \Psi_k | \Psi_l \rangle = \delta_{kl} \quad (7)$$

The Schrodinger Equation

Expectation values of observables are given by:

$$\langle \hat{A} \rangle = \frac{\int \Psi^* \hat{A} \Psi dx}{\int \Psi^* \Psi dx} = \frac{\langle \Psi | \hat{A} | \Psi \rangle}{\langle \Psi | \Psi \rangle} \quad (8)$$

In particular we also have

$$T[\Psi] = \langle \hat{T} \rangle = \int \Psi^* \hat{T} \Psi dx \quad (9)$$

and

$$V[\Psi] = \langle \hat{V} \rangle = \int \Psi^* \hat{V} \Psi dx \quad (10)$$

Variational Principle

For a system in a state Ψ

$$E[\Psi] = \frac{\langle \Psi | \hat{H} | \Psi \rangle}{\langle \Psi | \Psi \rangle} \quad (11)$$

where $\langle \Psi | \hat{H} | \Psi \rangle = \int \Psi^* \hat{H} \Psi dx$ The variational principle states that

$$E[\Psi] \geq E_0 \quad (12)$$

The total energy computed for a guessed state Ψ is an upper bound to the true ground state energy E_0 .

$$E_0 = \min_{\Psi} E[\Psi]. \quad (13)$$

Variational Principle

Formal proof of the variational principle.

Let's expand Ψ in terms of the normalized eigenstates of \hat{H} , Ψ_k :

$$\Psi = \sum_k C_k \Psi_k \quad (14)$$

The the energy becomes:

$$E[\Psi] = \frac{\sum_k |C_k|^2 E_k}{\sum_k |C_k|^2} \geq \frac{\sum_k |C_k|^2 E_0}{\sum_k |C_k|^2} \geq E_0 \quad (15)$$

since $E_k \geq E_0$.

(E_k is the energy for the k th eigenstate of \hat{H}).

Variational Principle

The variational principle can also be state in the variational form:

$$\delta \left[\langle \Psi | \hat{H} | \Psi \rangle - E \langle \Psi | \Psi \rangle \right] = 0 \quad (16)$$

where E is the Lagrange multiplier.

This eq guarantees extremization of $\langle \Psi | \hat{H} | \Psi \rangle$ under the constraint $\langle \Psi | \Psi \rangle = 1$.

The variational principle can be extended to excited states

$$E \left[\tilde{\Psi}_1 \right] \geq E_1 \quad (17)$$

if $\tilde{\Psi}_1$ is orthogonal to Ψ_0 .

The Hartree Fock Approximation

$$\Psi_{HF} = \frac{1}{\sqrt{N!}} \begin{vmatrix} \psi_1(x_1) & \psi_2(x_1) & \dots & \psi_N(x_1) \\ \psi_1(x_2) & \psi_2(x_2) & \dots & \psi_N(x_2) \\ \cdot & \cdot & & \cdot \\ \cdot & \cdot & & \cdot \\ \cdot & \cdot & & \cdot \\ \psi_1(x_N) & \psi_2(x_N) & \dots & \psi_N(x_N) \end{vmatrix} = \frac{1}{\sqrt{N!}} \det[\psi_1 \psi_2 \dots \psi_N] \quad (18)$$

orthogonal orbitals ψ_i are found minimizing $E[\Psi] = \frac{\langle \Psi | \hat{H} | \Psi \rangle}{\langle \Psi | \Psi \rangle}$ where Ψ is in the form of a Slater determinant. $\langle \Psi_{HF} | \Psi_{HF} \rangle = 1$

The Hartree Fock Approximation

$$E_{HF} = \langle \Psi_{HF} | \hat{H} | \Psi_{HF} \rangle = \sum_{i=1}^N H_i + \frac{1}{2} \sum_{i,j=1}^N (J_{ij} - K_{ij}) \quad (19)$$

where

$$H_i = \int \psi_i^*(\mathbf{x}) \left[-\frac{1}{2} \nabla^2 + v(\mathbf{x}) \right] \psi_i(\mathbf{x}) d\mathbf{x} \quad (20)$$

$$J_{ij} = \iint \psi_i(\mathbf{x}_1) \psi_i^*(\mathbf{x}_1) \frac{1}{r_{12}} \psi_j^*(\mathbf{x}_2) \psi_j(\mathbf{x}_2) d\mathbf{x}_1 d\mathbf{x}_2 \quad (21)$$

$$K_{ij} = \iint \psi_i^*(\mathbf{x}_1) \psi_j(\mathbf{x}_1) \frac{1}{r_{12}} \psi_i(\mathbf{x}_2) \psi_j^*(\mathbf{x}_2) d\mathbf{x}_1 d\mathbf{x}_2 \quad (22)$$

$$J_{ii} = K_{ii} \quad (23)$$

this explain the double sum in (19).

The Hartree Fock Approximation

Minimizing E_{HF}

$$E_{HF} = \langle \Psi_{HF} | \hat{H} | \Psi_{HF} \rangle = \sum_{i=1}^N H_i + \frac{1}{2} \sum_{i,j=1}^N (J_{ij} - K_{ij}) \quad (24)$$

subject to the orthonormalization conditions

$$\int \psi_i^*(x) \psi_j(x) dx = \delta_{ij} \quad (25)$$

gives the Hartree Fock differential equations:

$$\hat{F} \psi_i(x) = \sum_{j=1}^N \epsilon_{ij} \psi_j(x) \quad (26)$$

$\hat{F}\psi_i(x) = \sum_{k=1}^N \epsilon_{ij}\psi_j(x)$ where

$$\hat{F} = -\frac{1}{2}\nabla^2 + v(x) + \hat{g} \quad (27)$$

in which the Coulomb-exchange operator \hat{g} is given by

$$\hat{g} = \hat{j} - \hat{k}. \quad (28)$$

Here

$$\hat{j}(x_1)f(x_1) \equiv \sum_{k=1}^N \int \psi_k^*(x_2)\psi_k(x_2) \frac{1}{r_{12}} f(x_1) dx_2 \quad (29)$$

and

$$\hat{k}(x_1)f(x_1) \equiv \sum_{k=1}^N \int \psi_k^*(x_2)f(x_2) \frac{1}{r_{12}} \psi_k(x_1) dx_2 \quad (30)$$

where $f(x_1)$ is an arbitrary function and the matrix ϵ consists of Lagrange multipliers.

Restricted Hartree Fock (RHF)

For an even number of electrons the alpha and beta electron are chosen so occupy the same orbitals, namely we have $N/2$ orbitals of form $\phi_k(r)\alpha(s)$ and $N/2$ orbitals of form $\phi_k(r)\beta(s)$

$$E_{HF} = 2 \sum_{k=1}^{N/2} H_k + \sum_{k,l=1}^{N/2} (2J_{kl} - K_{kl}) \quad (31)$$

where

$$H_k = \int \phi_k^*(r) \left[-\frac{1}{2} \nabla^2 + v(r) \right] \phi_k(r) dr \quad (32)$$

$$J_{kl} = \iint \phi_k(r_1) \phi_k^*(r_1) \frac{1}{r_{12}} \phi_l^*(r_2) \phi_l(r_2) dr_1 dr_2 \quad (33)$$

$$K_{kl} = \iint \psi_k^*(r_1) \psi_l(r_1) \frac{1}{r_{12}} \psi_k(r_2) \psi_l^*(r_2) dr_1 dr_2 \quad (34)$$

The Hartree-Fock equations now read:

$$\hat{F}\phi_k(r) = \sum_{l=1}^{N/2} \epsilon_{kl}\phi_l(r) \quad (35)$$

with operator $\hat{F} = -\frac{1}{2}\nabla^2 + v(x) + \hat{g}$ ($\hat{g} = \hat{j} - \hat{k}$) where

$$\hat{j}(r_1)f(r_1) \equiv 2 \sum_{m=1}^{N/2} \int \phi_m^*(r_2)\phi_m(r_2) \frac{1}{r_{12}} dr_2 f(r_1) \quad (36)$$

and

$$\hat{k}(r_1)f(r_1) \equiv \sum_{m=1}^{N/2} \int \phi_m^*(r_2)f(r_2) \frac{1}{r_{12}} dr_2 \phi_m(r_1) dr_2 \quad (37)$$

the determinant wavefunction for the "closed-shell" case is

$$\Psi_{HF} = \frac{1}{\sqrt{N!}} \begin{vmatrix} \phi_1(r_1)\alpha(s_1) & \phi_1(r_1)\beta(s_1) & \dots & \phi_{N/2}(r_1)\beta(s_1) \\ \phi_1(r_2)\alpha(s_2) & \phi_1(r_2)\beta(s_2) & \dots & \phi_{N/2}(r_2)\beta(s_2) \\ \cdot & \cdot & & \cdot \\ \cdot & \cdot & & \cdot \\ \cdot & \cdot & & \cdot \\ \phi_1(r_N)\alpha(s_N) & \phi_1(r_N)\beta(s_N) & \dots & \phi_{N/2}(r_N)\beta(s_N) \end{vmatrix} \quad (38)$$

A unitary transformation of the occupied orbitals ϕ_k to another set of orbitals η_k leaves the wavefunction unchanged except possibly by a phase factor. Also the operators \hat{F} , \hat{k} and \hat{j} are invariant to such a transformation.

If we have

$$\eta_m = \sum_k U_{mk} \psi_k \quad (39)$$

where

$$U^+ U = 1 \quad (40)$$

then the Hartree Fock equations become:

$$\hat{F} \eta_m = \sum_{n=1}^{N/2} \epsilon_{mn}^{\eta} \eta_n \quad (41)$$

One can choose U so to diagonalize the Hartree Fock equations:

$$\hat{F} \lambda_m(r) = \epsilon_m^{\lambda} \lambda_m(r) \quad (42)$$

The orbitals solution of

$$\hat{F}\lambda_m(r) = \epsilon_m^\lambda \lambda_m(r)$$

are uniquely appropriate for describing removal of electrons from the system.

Koopmans theorem If one assume no reorganization on ionization, then the best (lowest energy) single-determinant description for the ion is the determinant built from the canonical Hartree Fock orbitals $\lambda_m(r)$.

$$\epsilon_m^\lambda = -I_m \quad (43)$$

where I_m is the ionization energy associated with the removal of an electron from the orbital λ_m .

When the number of electron is not even, the standard HF approach is called *unrestricted open-shell* HF method.

spatial part of wf for spin α is allowed to be different from the spatial part of the wf for spin β .

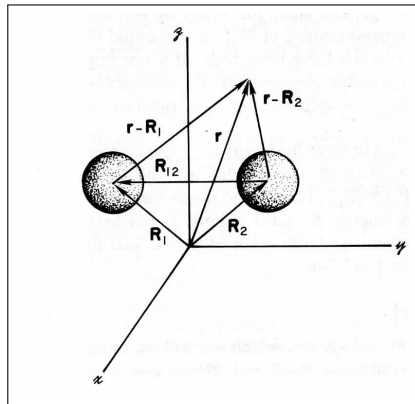
UHF can be used also in the case of even number of electrons. It often doesn't give a lower energy, but there are important examples where the UHF is required. (e.g. H_2).

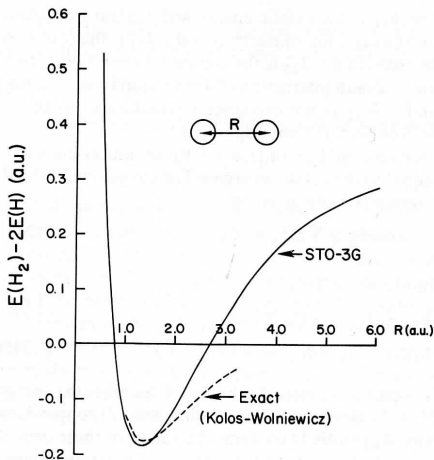
UHF for H_2 molecule

$$\phi_{1s}^{CGF} (\zeta = 1.0, STO - 3G) = d_{13}\phi_{1s}^{GF}(\alpha_{13}) + d_{23}\phi_{1s}^{GF}(\alpha_{23}) + d_{33}\phi_{1s}^{GF}(\alpha_{33})$$

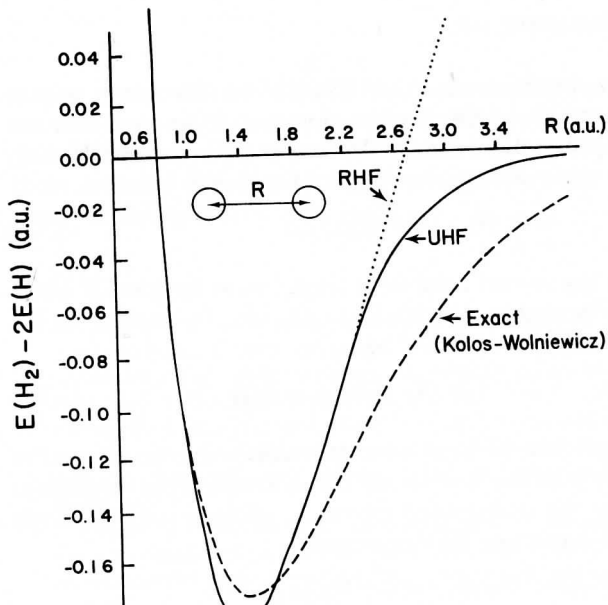
where

$$\phi_{1s}^{GF}(\alpha, r - R_A) = (2\alpha/\pi)^{3/4} e^{-\alpha|r-R_A|^2}$$





Kolos-Wolniewicz, J. Chem. Phys. 41, 3663 (1964); doi:10.1063/1.1725796



Correlation energy

HF is a single determinant description, but the Exact solution is never a single determinant or a combination of a few determinants. The energy difference between the exact energy and HF energy is the correlation energy.

$$E_{corr} = E - E_{HF} < 0 \quad (44)$$

How to improve?

- ▶ CI (linear mixing of several determinants. Conceptually the simplest, but NOT computationally

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How to improve?

- ▶ CI (linear mixing of several determinants. Conceptually the simplest, but NOT computationally)
- ▶ Many body perturbation theory

Electron Density

$$\rho(r_1) = N \int \dots \int |\Psi(x_1, x_2, \dots, x_N)|^2 ds_1 dx_2 \dots dx_N \quad (45)$$

$$\int \rho(r) dr = N \quad (46)$$

- ▶ For an atom in its ground state, the density decrease monotonically away from the nucleus
- ▶ At any atomic nucleus, the electron density has a finite value. Cusp in the density to avoid $-\frac{1}{2}\nabla^2 - (Z_\alpha/r_\alpha)$ to blow up in $\hat{H}\Psi$.

$$\frac{\partial}{\partial r_\alpha} \bar{\rho}(r_\alpha)|_{\alpha=0} = -2Z_\alpha \bar{\rho}(0) \quad (47)$$

where $\bar{\rho}(r_\alpha)$ is the spherical average of $\rho(r_\alpha)$.

Electron Density

Long-range law for the electron density:

$$\rho \sim \exp \left[-2 (2I_{min})^{1/2} r \right] \quad (48)$$

where I_{min} is the first ionization potential.

The Hartree Fock result following from Koopmans theorem will be:

$$\rho_{HF} \sim \exp \left[-2 (-2\epsilon_{max})^{1/2} r \right] \quad (49)$$

The original idea: the Thomas-Fermi model

The fundamental idea is to replace the N-electron wavefunction with the electron density.

$$\psi(x_1, x_2, \dots, x_N) \rightarrow \rho(r) \quad (50)$$

How to approximate the the distribution of electrons in an atom?
Let's assume electrons are uniformly distributed

$$\epsilon(n_x, n_y, n_z) = \frac{h^2}{8ml^2} (n_x^2 + n_y^2 + n_z^2) \quad (51)$$

$$\epsilon(n_x, n_y, n_z) = \frac{h^2}{8ml^2} R^2 \quad (52)$$

number of distinct energy levels with energy lower than ϵ

$$\Phi(\epsilon) = \frac{1}{8} \frac{4\pi}{3} R^3 \quad (53)$$

$$\Phi(\epsilon) = \frac{\pi}{6} \left(\frac{8mI^2}{h^2} \right)^{3/2} \epsilon^{3/2} \quad (54)$$

The number of energy levels between ϵ and $\epsilon + \delta\epsilon$ is

$$g(\epsilon)\Delta\epsilon = \Phi(\epsilon + \delta\epsilon) - \Phi(\epsilon) \quad (55)$$

$$= \frac{\pi}{4} \left(\frac{8mI^2}{h^2} \right)^{3/2} \epsilon^{1/2} \delta\epsilon \quad (56)$$

$$\Delta E = 2 \int \epsilon f(\epsilon) g(\epsilon) d\epsilon \quad (57)$$

$$= 2 \int_0^{\epsilon_F} \frac{\pi}{4} \left(\frac{8ml^2}{h^2} \right)^{3/2} \epsilon^{3/2} \delta\epsilon \quad (58)$$

$$= \frac{\pi}{5} \left(\frac{8ml^2}{h^2} \right)^{3/2} \epsilon_F^{5/2} \quad (59)$$

$$= \frac{8\pi}{5} \left(\frac{2m}{h^2} \right)^{3/2} l^3 \epsilon_F^{5/2} \quad (60)$$

where we have used the Fermi-Dirac distribution for the electrons:

$$f(\epsilon) = \frac{1}{1 + e^{\beta(\epsilon - \mu)}} \quad (61)$$

which at $T=0$ becomes

$$f(\epsilon) = \begin{cases} 1 & \text{if } \epsilon < \epsilon_F \\ 0 & \text{if } \epsilon > \epsilon_F \end{cases}$$

$$\Delta N = 2 \int f(\epsilon)g(\epsilon)d\epsilon \quad (62)$$

$$= \frac{8\pi}{3} \left(\frac{2m}{h^2} \right)^{3/2} I^3 \epsilon_F^{3/2} \quad (63)$$

$$\Delta E = \frac{3}{5} \Delta N \epsilon_F \quad (64)$$

$$= \frac{3h^2}{10m} \left(\frac{3}{8\pi} \right)^{2/3} \left(\frac{\Delta N}{I^3} \right)^{5/3} \quad (65)$$

$$= \frac{3h^2}{10m} \left(\frac{3}{8\pi} \right)^{2/3} \rho^{5/3} \quad (66)$$

Adding the contributions from all the cells we get

$$T_{TF}[\rho] = C_F \int \rho^{5/3}(r) dr \quad (67)$$

$$T_{TF}[\rho] = C_F \int \rho^{5/3}(r) dr \quad (68)$$

Here we have reverted to atomic units

Fundamental atomic units

Dimension	Name	Symbol/Definition	Value in SI units ^[5]
mass	electron rest mass	m_e	$9.109\,382\,91(40) \times 10^{-31}$ kg
charge	elementary charge	e	$1.602\,176\,565(35) \times 10^{-19}$ C
action	reduced Planck's constant	$\hbar = h/(2\pi)$	$1.054\,571\,726(47) \times 10^{-34}$ J·s
electric constant ⁻¹	Coulomb force constant	$k_e = 1/(4\pi\epsilon_0)$	$8.987\,551\,787\,3681 \times 10^9$ kg·m ³ ·s ⁻² ·C ⁻²

and introduced the constant:

$$C_F = \frac{3}{10} (3\pi^2)^{2/3} = 2.871 \quad (69)$$

If we now come back to the Hamiltonian for an atomic system

$$\hat{H} = \sum_{i=1}^N \left(-\frac{1}{2} \nabla_i^2\right) + \sum_{i=1}^N v(r_i) + \sum_{i<j}^N \frac{1}{r_{ij}} \quad (70)$$

we can write the Thomas-Fermi functional of atoms

$$E_{TF}[\rho] = C_F \int \rho^{5/3}(r) dr - Z \int \frac{\rho(r)}{r} dr + \frac{1}{2} \int \int \frac{\rho(r_1)\rho(r_2)}{|r_1 - r_2|} dr_1 dr_2 \quad (71)$$

We now assume that for the ground state the electron density minimizes the energy functional under the condition:

$$N = N[\rho(r)] = \int \rho(r) dr \quad (72)$$

Or in the Lagrange multiplier formalism:

$$\delta\{E_{TF}[\rho] - \mu_{TF} \left(\int \rho(r) dr - N \right)\} = 0 \quad (73)$$

which yields the Euler-Lagrange equation

$$\mu_{TF} = \frac{\delta E_{TF}[\rho]}{\delta \rho(r)} = \frac{5}{3} C_F \rho^{2/3}(r) - \phi(r) \quad (74)$$

where $\phi(r)$ is the electrostatic potential

$$\phi(r) = \frac{Z}{r} - \int \frac{\rho(r_2)}{|r - r_2|} dr_2 \quad (75)$$

At the beginning the model encountered limited success, indeed

- ▶ Accuracy for atoms is not high
- ▶ No molecular binding is predicted (see Teller, E. (1962). "On the Stability of molecules in the Thomas-Fermi theory". Rev. Mod. Phys. 34 (4): 627-631.

However the situation changed in 1964 with the publication of a landmark paper by Hohenberg and Kohn. They provided the fundamental theorems to show that:

- ▶ Thomas-Fermi model can be regarded as an approximation to an exact theory, the *Density Functional Theory*.

The Hohenberg-Kohn theorems

For a given system described by the Hamiltonian:

$$\hat{H} = \sum_{i=1,}^N \left(-\frac{1}{2} \nabla_i^2\right) + \sum_{i=1,}^N v(r_i) + \sum_{i<j}^N \frac{1}{r_{ij}} \quad (76)$$

N and $v(r)$ determines all properties of the ground state. In place of N and $v(r)$, the **first Hohenberg-Kohn theorem** legitimizes the use of the electron density $\rho(r)$ as basic variable. It states:

The external potential $v(r)$ is determined, within a trivial additive constant, by the electron density $\rho(r)$.

Let's consider the ground state density $\rho(r)$.

$\rho(r)$ determines N as $N = \int \rho(r) dr$.

$\rho(r)$ also determines $v(r)$. Let's say that is not true and there are two distinct $v(r)$ and $v'(r)$, differing by more than a constant, both giving the same $\rho(r)$. We would have two Hamiltonian H and H' . Using the variational principle we can write

$$E_0 < \langle \psi' | H | \psi' \rangle = \langle \psi' | H - H' | \psi' \rangle + \langle \psi' | H' | \psi' \rangle \quad (77)$$

$$= \int \rho(r) [v(r) - v'(r)] dr + E'_0 \quad (78)$$

In the same way we can write

$$E'_0 < \langle \psi | H' | \psi \rangle = \langle \psi | H' - H | \psi \rangle + \langle \psi | H | \psi \rangle \quad (79)$$

$$= - \int \rho(r) [v(r) - v'(r)] dr + E_0 \quad (80)$$

Summing up the two we arrive to the absurd $E_0 + E'_0 < E'_0 + E_0$, so it must be $v(r) = v'(r)$.

The **second Hohenberg-Kohn theorem** states:

For a trial density $\tilde{\rho}(r)$ such that $\tilde{\rho}(r) > 0$ and $\int \tilde{\rho}(r) dr = N$, there is a variational principle and $E_0 \leq E_v[\tilde{\rho}]$.

$$E_v[\rho] = T[\rho] + V_{ne}[\rho] + V_{ee}[\rho] \quad (81)$$

$$= \int \rho(r) v(r) dr + F_{HK}[\rho] \quad (82)$$

and

$$F_{HK}[\rho] = T[\rho] + V_{ee}[\rho] \quad (83)$$

$$= T[\rho] + J[\rho] + \text{nonclassical term} \quad (84)$$

The second HK theorem provides the justification for the variational principle in TF, where $E_{TF}[\rho]$ is an approximation to $E[\rho]$.

From the first HK Theorem we have that $\tilde{\rho}(r)$ determines $\tilde{v}(r)$ and consequently \tilde{H} and $\tilde{\psi}$.
Let's consider $\tilde{\psi}$ as the trial wf for the problem H

$$\langle \tilde{\psi} | H | \tilde{\psi} \rangle = \int \tilde{\rho}(r) dr + F_{HK}[\tilde{\rho}] = E_v[\tilde{\rho}] \quad (85)$$

$$\langle \tilde{\psi} | H | \tilde{\psi} \rangle \geq \langle \psi | H | \psi \rangle = E_v[\rho] \quad (86)$$

So it follows that $E_v[\tilde{\rho}] \geq E_v[\rho]$.

In the differential form:

$$\delta \left\{ E_v[\rho] - \mu \left[\int \rho(r) dr - N \right] \right\} = 0 \quad (87)$$

where

$$\mu = \frac{\delta E_v[\rho]}{\delta \rho(r)} = v(r) + \frac{\delta F_{HK}[\rho]}{\delta \rho(r)} \quad (88)$$

Orbitals and Kohn-Sham equations

it is quite appealing that we can get the ground-state energy of a many-electron system as minimum of an energy functional

$$E_v[\rho] = T[\rho] + V_{ne}[\rho] + V_{ee}[\rho]$$

we have seen that a drastic approximation to the functional can be obtained by the TF model:

$$V_{ee}[\rho] \rightarrow J[\rho] = \frac{1}{2} \iint \frac{\rho(r_1)\rho(r_2)}{|r_2 - r_1|} dr_1 dr_2$$

$$T[\rho] \rightarrow \text{uniform electron gas}$$

- ▶ Can we do better?

Let's consider a *noninteracting reference system* in which there is no electron electron repulsion term and for which the electron density is exactly ρ :

$$\hat{H}_s = \sum_i^N \left(-\frac{1}{2} \nabla_i^2 \right) + \sum_i^N v_s(r_i) \quad (89)$$

For this system we have an exact solution

$$\Psi_s = \frac{1}{\sqrt{N!}} \det [\psi_1 \psi_2 \dots \psi_N] \quad (90)$$

for such a system the kinetic energy is

$$T_s[\rho] = \langle \Psi_s | \sum_i^N \left(-\frac{1}{2} \nabla_i^2 \right) | \Psi_s \rangle = \quad (91)$$

$$= \sum_i^N \langle \Psi_s | -\frac{1}{2} \nabla_i^2 | \Psi_s \rangle \quad (92)$$

Now the very clever idea by Kohn and Sham was to rewrite

$$E_v[\rho] = T[\rho] + V_{ne}[\rho] + V_{ee}[\rho]$$

as

$$E_v[\rho] = T_s[\rho] + V_{ne}[\rho] + J[\rho] + E_{xc}[\rho] \quad (93)$$

where

$$E_{xc}[\rho] = T[\rho] - T_s[\rho] + V_{ee}[\rho] - J[\rho] \quad (94)$$

The quantity $E_{xc}[\rho]$ is the *exchange-correlation energy*, which contains the difference between $T[\rho]$ and $T_s[\rho]$ and the non-classical part of $V_{ee}[\rho]$.