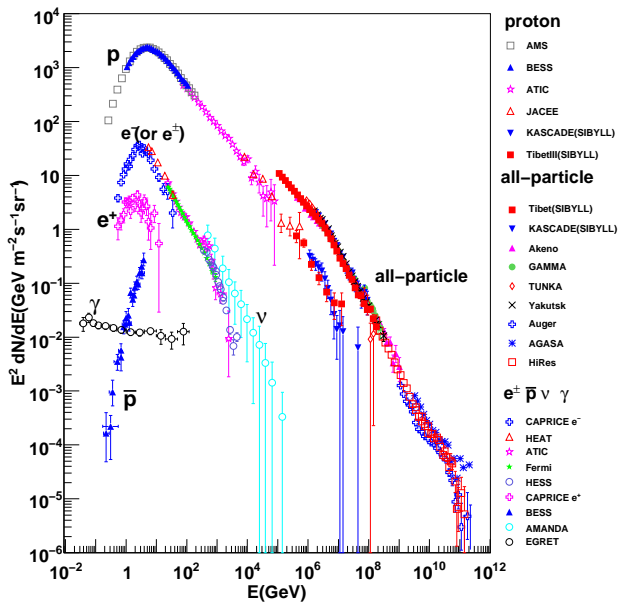


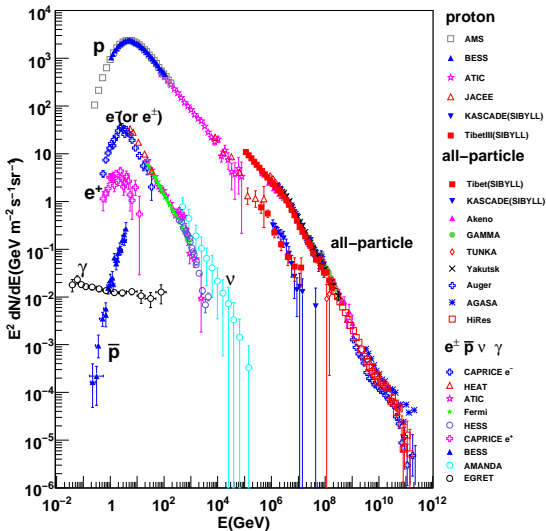
Cosmic Ray Introduction

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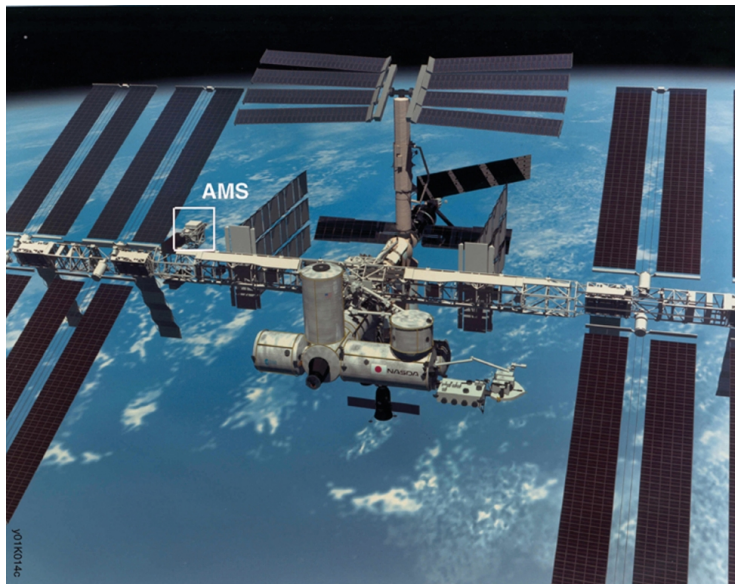


- Flux at 100 GeV:
~ 1 particle/m²/sec
- Flux at 10⁶ GeV (“knee”):
~ 1 particle/m²/year
- Flux at 10¹⁰ GeV (“ankle”):
~ 1 particle/km²/year
- At 10¹¹ GeV: endpoint
 (“GZK cutoff”, see later)

Tools for studying cosmic rays

- Space-based experiments
 - ▶ PAMELA, AMS-02, Fermi, ...
 - ▶ $E \lesssim 10^3$ GeV
 - ▶ Clean measurement, long observation times

AMS-02 (Alpha Magnetic Spectrometer)



Tools for studying cosmic rays

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- Balloon experiments
 - ▶ $E \lesssim 10^6$ GeV
 - ▶ Only short observation periods
 - ▶ Atmospheric distortions
- Imaging Air Čerenkov Telescopes
 - ▶ HESS, MAGIC, VERITAS, ...
 - ▶ 10 GeV $\lesssim E \lesssim 10^6$ GeV

H.E.S.S. (High Energy Stereoscopic System)



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- Imaging Air Čerenkov Telescopes
 - ▶ HESS, MAGIC, VERITAS, ...
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- Air shower detectors
 - ▶ Pierre Auger observatory, ...
 - ▶ 10^5 GeV $\lesssim E \lesssim 10^{12}$ GeV

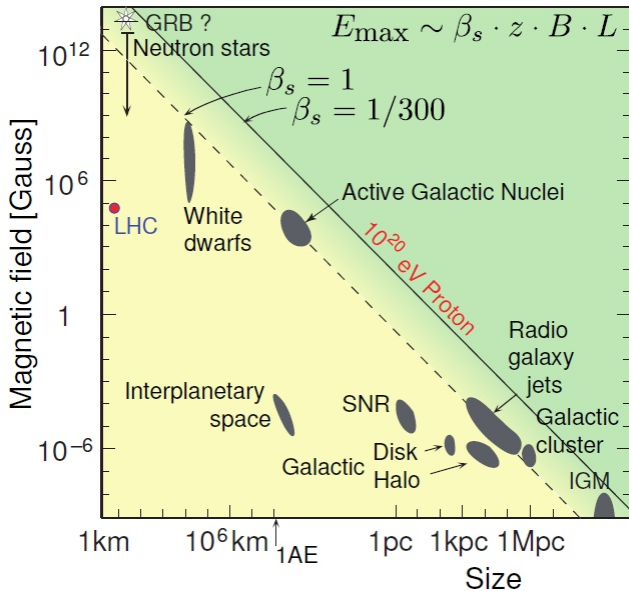
Pierre Auger Observatory



Pierre Auger Observatory



Hillas Plot



Pulsars

