

Neutrino physics

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1. Neutrinos and their masses

1.1 The role of neutrinos in physics

The standard model fermions

$$\text{Quarks} \begin{cases} u & c & t & q = +\frac{2}{3} \\ d & s & b & q = -\frac{1}{3} \end{cases}$$

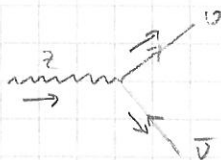
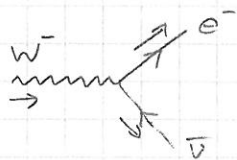
$$\text{Leptons} \begin{cases} \nu_e & \nu_\mu & \nu_\tau & q = 0 \\ e & \mu & \tau & q = -1 \end{cases}$$

Role in physics

- Produced in nuclear reactions, e.g. $(A, Z) \rightarrow (A, Z+1) + e^- + \bar{\nu}_e$
- $\pi^+ \rightarrow \mu^+ + \nu_\mu$, $p + p \rightarrow D + e^+ + \nu_e$
- May help elucidate the origin of flavor (why 3 families? why small mixing in quark sector, large in neutrino sector?)
- May be sensitive to GUT scale physics
- May help to explain the baryon asymmetry of the Universe
- can teach us about dark matter
- can teach us about sources of high-E cosmic rays
- ...

Lagrangian:

$$\mathcal{L} = \sum_{\alpha=e,\mu,\tau} \left[\bar{\nu}_{\alpha,L} i \not{\partial} \nu_{\alpha,L} + \frac{g}{\sqrt{2}} (W^{\mu+} \bar{\nu}_{\alpha,L} \gamma_\mu e_{\alpha,L} + \text{h.c.}) + \frac{g}{2 \cos \theta_w} Z^\mu \bar{\nu}_{\alpha,L} \gamma_\mu \nu_{\alpha,L} \right] + \text{mass term}$$



1.2 Dirac mass terms

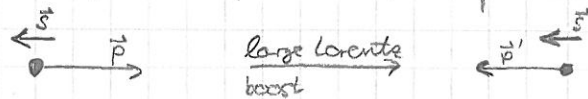
α involves only $\psi_L = \frac{1-\gamma^5}{2} \psi = \begin{pmatrix} \chi_1 \\ \chi_2 \\ 0 \\ 0 \end{pmatrix}$

... only upper components of 4-component spinor are physical
(LH $\nu + \bar{\nu}$)

(Dirac) mass terms make ψ_R physical:

$$\mathcal{L}_m \supset \sum_{\alpha, \beta} m_{\nu, \alpha \beta} \bar{\psi}_{L, \alpha} \psi_{R, \beta} + \text{h.c.}$$

Physical reason for 4 d.o.f.: Consider particle with LH helicity:



\Rightarrow states with RH helicity must exist \Rightarrow 2 d.o.f.; + antiparticles: 4 d.o.f

Problem: Why is m_ν so much smaller than other fermion masses?

1.3 Majorana mass terms

Mass terms couple LH and RH fields.

The antiparticle of ν_L is a RH particle.

Could ν_R be identical to the antiparticle of ν_L ?

[Note: This can work only for neutrinos — for all other SM fermions, antiparticles carry opposite charge \Rightarrow have to be different d.o.f.]

More formally: charge conjugation:

$$\hat{C}: \psi \rightarrow \psi^c \equiv \underbrace{-i\gamma^2\gamma^0}_{\equiv C} \bar{\psi}^T = -i\gamma^2\psi^*$$

Effect on chirality:

$$\gamma^5 \psi^c = +i\gamma^2 \gamma^5 \psi^* = -(\gamma^5 \psi)^c$$

$\Rightarrow \hat{C}$ flips chirality, transforms LH particle \leftrightarrow RH antiparticle

Identify $\nu_R \equiv (\nu_L)^c$

In 4-component spinor notation: $\nu_L = \begin{pmatrix} \chi \\ 0 \end{pmatrix}$

$$(\nu_L)^c = -i \begin{pmatrix} 0 & \sigma^2 \\ -\sigma^2 & 0 \end{pmatrix} \begin{pmatrix} \chi^* \\ 0 \end{pmatrix} = \begin{pmatrix} 0 \\ i\sigma^2 \chi^* \end{pmatrix}$$

$$\Rightarrow \nu = \begin{pmatrix} \chi \\ i\sigma^2 \chi^* \end{pmatrix}; \quad \nu^c = \nu$$

only 2 dof!

A new type of mass term:

$$\mathcal{L}_m \supseteq \sum_{\alpha\beta} m_{\alpha\beta} \overline{(\nu_L)^c}_{\alpha} \nu_{L\beta} + \text{h.c.}$$

Problems: - How to obtain from SU(2)-invariant theory?
- Why is m so small?

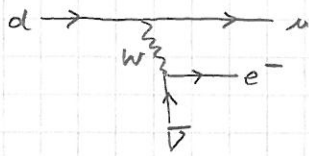
1.4 Measuring neutrino mass

1.4.1 Kinematic measurement

β -decay, e.g. ${}^3\text{H} \rightarrow {}^3\text{He} + e^- + \bar{\nu}_e$

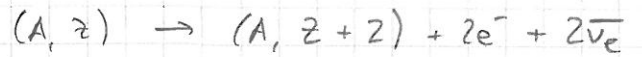
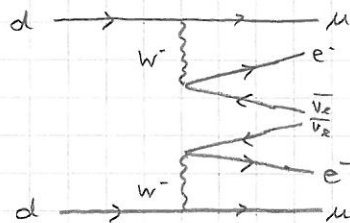
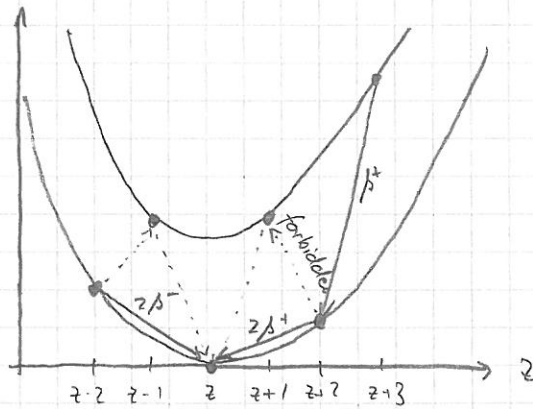
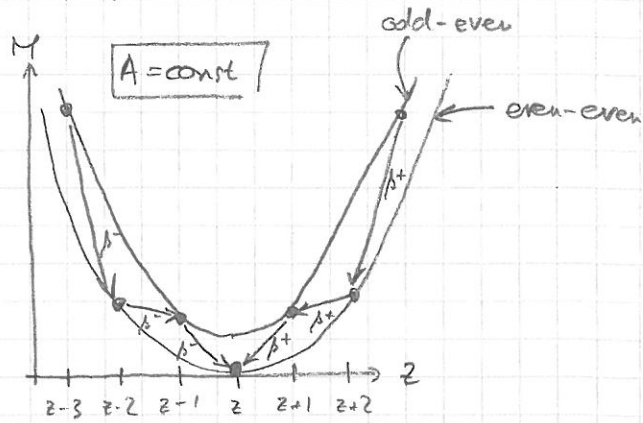
$$E_{e,\text{max}} = \underbrace{Q}_{m_{\text{H}} - m_{\text{He}}} - m_{\nu}$$

\Rightarrow measure e^- spectrum precisely

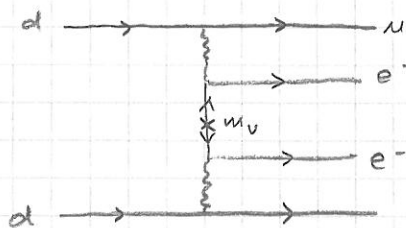


1.4.2 Neutrinoless double beta decay

Nuclear masses:



For Majorana neutrinos, also allowed



1.4.3 Cosmology

→ lecture by J. Jedam

neutrinos transport energy efficiently over large distances → wash out structure on small scales

1.5 The seesaw mechanism

Goal: Explain why neutrino masses are so small

Augment SM with 3 RH neutrinos N_{Rj} , singlet under SM gauge group.

(Here: Consider just 1H neutrino + 1 RH neutrino for simplicity)

$$\mathcal{L} = - \underbrace{m_D}_{O(m_H)} \bar{\nu}_L N_R + \frac{1}{2} \underbrace{m_H}_{\text{can be very large}} \overline{(N_R)^c} N_R + \text{h.c.} \quad (*)$$

↑
see exercise

Write $n \equiv \begin{pmatrix} \nu_L \\ (N_R)^c \end{pmatrix}$

$$\hookrightarrow \mathcal{L} = -\frac{1}{2} \bar{n}^c M n + \text{h.c.}$$

$$= -\frac{1}{2} \begin{pmatrix} \overline{(\nu_L)^c} & \overline{N_R} \end{pmatrix} \begin{pmatrix} 0 & m_D \\ m_D & m_H \end{pmatrix} \begin{pmatrix} \nu_L \\ (N_R)^c \end{pmatrix} + \text{h.c.} \quad (**)$$

(see exercise for proof that (*) and (**) are equivalent.)

Diagonalize $M \rightarrow$ eigenvalues $\sim -\frac{m_D^2}{m_H}, m_H$

eigenvectors $\nu_L' = \nu_L + O\left(\frac{m_D}{m_H}\right) (N_R)^c$

$N_R' = N_R + O\left(\frac{m_D}{m_H}\right) (\nu_L)^c$

Transform $\nu_L' \rightarrow i \nu_L'$ to absorb negative sign of eigenvalue into redefinition of phase of field

$$\hookrightarrow \mathcal{L} = -\frac{1}{2} \underbrace{\frac{m_D^2}{m_H}}_{\equiv m_\nu} \overline{(\nu_L')^c} \nu_L'$$

With $m_D \sim 100 \text{ GeV}$, $m_H \sim 10^{14} \text{ GeV}$: $m_\nu = \frac{m_D^2}{m_H} \sim 0.1 \text{ eV}$

2. Neutrino mixing and oscillations

2.1 Three flavors of neutrinos

Remember:

$$\mathcal{L} \supset \sum_{\alpha=e,\mu,\tau} \left[\bar{\nu}_{\alpha,L} i \not{\partial} \nu_{\alpha,L} + \frac{g}{\sqrt{2}} (W^{\mu+} \bar{\nu}_{\alpha,L} \gamma_{\mu} e_{\alpha,L} + \text{h.c.}) \right. \\ \left. + \frac{g}{2\cos\theta_W} Z^{\mu} \bar{\nu}_{\alpha,L} \gamma_{\mu} \nu_{\alpha,L} \right] \\ + \left(\sum_{\alpha,\beta} \frac{1}{2} m_{\alpha\beta} \overline{(\nu_{\alpha})^c} \nu_{\beta} + \text{h.c.} \right)$$

- $m_{\alpha\beta}$ in general off-diagonal

- Diagonalization: Write $\nu_{L\alpha} = U_{\alpha j} \nu_j$
with $U^T m U = \text{diag}(m_1, m_2, m_3)$

For this to work, m must be symmetric, $m_{\alpha\beta} = m_{\beta\alpha}$

- Note. For Dirac mass-term $\sum_{\alpha\beta} m_{\alpha\beta} \bar{\nu}_{L\alpha} \nu_{R\beta}$,

the transformation is $\nu_{L\alpha} = U_{\alpha j} \nu_j$; $\nu_{R\beta} = V_{\beta j} \nu_{Rj}$

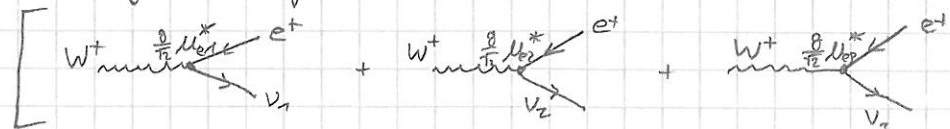
$$U^T m V = \text{diag}(m_1, m_2, m_3)$$

This works for arbitrary complex m .

- In the mass basis:

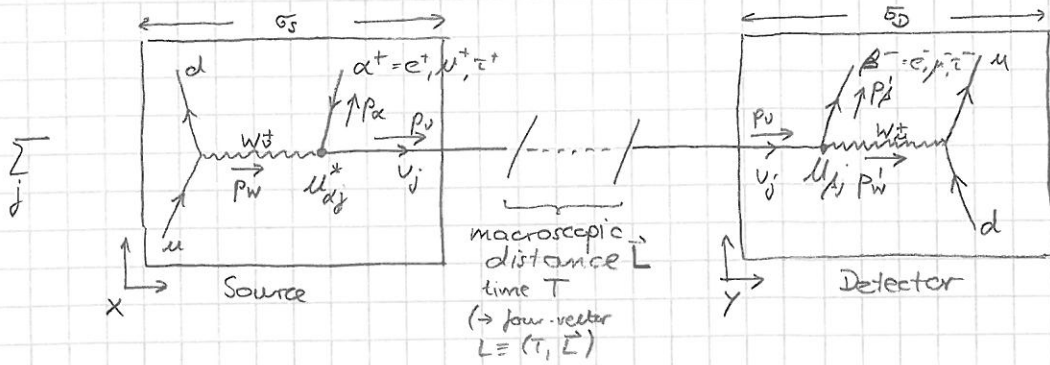
$$\mathcal{L} \supset \bar{\nu}_{Lj} i \not{\partial} \nu_{Lj} + \frac{g}{\sqrt{2}} (W^{\mu+} \bar{\nu}_{Lj} U_{\alpha j}^* \gamma_{\mu} e_{\alpha} + \text{h.c.}) \\ + \frac{g}{2\cos\theta_W} Z^{\mu} \bar{\nu}_{Lj} \gamma_{\mu} \nu_{Lj} + \left(\frac{1}{2} m_j \overline{(\nu_{Lj})^c} \nu_{Lj} + \text{h.c.} \right)$$

\Rightarrow CC neutrino interactions produce superposition of mass eigenstates ν_j :



2.2 Neutrino oscillations

Consider typical neutrino experiment

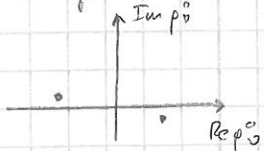


QFT

$$\begin{aligned}
 \mathcal{A} &\propto \int_{\text{Source}} d^4x \int_{\text{Detector}} d^4y \bar{u}_\beta(p'_\beta) e^{i p'_\beta y} \gamma^\mu U_{\beta k} \\
 &\cdot \int \frac{d^4 p_\nu}{(2\pi)^4} \frac{p_\nu e^{-i p_\nu (y+L-x)}}{p_\nu^2 - m_\nu^2 + i\epsilon} U_{\alpha j}^* \gamma^\nu U_{\alpha j} e^{-i p_\nu x} \\
 &\cdot \underbrace{M_\nu^S M_\mu^D}_{\text{hadronic matrix elements}} e^{i p_\nu y} e^{-i p_\nu x} \\
 &= \sum_j U_{\alpha j}^* U_{\beta j} M_\nu^S M_\mu^D \int \frac{d^4 p_\nu}{(2\pi)^4} \\
 &\cdot \hat{\delta}(p_\nu - p'_\beta - p'_\mu) \hat{\delta}(p_\nu - p_\alpha - p_\mu)
 \end{aligned}$$

approximate δ -functions, uncertainty $\sim (\text{size of } S, D)^{-1} \equiv \sigma_{S,D}^{-1}$

complex contour integration



$$\begin{aligned}
 &(p_0^2 - \sqrt{p^2 + m^2} + i\epsilon) \\
 &\cdot (p_0^2 - \sqrt{p^2 + m^2} - i\epsilon) \\
 &= (p_0^2 - (\sqrt{p^2 + m^2} - i\epsilon))^2 \\
 &= p_0^2 - m^2 + i\epsilon \\
 &= 2\sqrt{p^2 + m^2} \cdot \epsilon
 \end{aligned}$$

Consider phase $\phi(\vec{p}_\nu) \equiv E_\nu T - \vec{p}_\nu \vec{L}$

$$\phi(\vec{p}_\nu) \approx \vec{p}_\nu \left(T \frac{\vec{p}_\nu}{|\vec{p}_\nu|} - \vec{L} \right) + \frac{m_\nu^2}{2|\vec{p}_\nu|} T$$

QM

Produced neutrino state:

$$|\nu_\alpha\rangle = U_{\alpha j}^* | \nu_j \rangle$$

Detected neutrino state

$$\langle \nu_\beta | = U_{\beta k} \langle \nu_k |$$

Time and space evolution

$$|\nu_\alpha\rangle \rightarrow |\nu_\alpha(T, L)\rangle = \sum_j U_{\alpha j}^* e^{-i E_j T + i p_j L} | \nu_j \rangle$$

$$\mathcal{A} = \langle \nu_\beta | \nu_\alpha(T, L) \rangle$$

$$= \sum_{j,k} U_{\alpha j}^* U_{\beta k} e^{-i E_j T + i p_j L} \langle \nu_k | \nu_j \rangle$$

$$= \sum_j U_{\alpha j}^* U_{\beta k} e^{-i E_j T + i p_j L}$$

Note: States with different energies and momenta $(E_1, p_1), (E_2, p_2), (E_3, p_3)$ can interfere only if E, p -uncertainties in source, detector are larger than $|E_i - E_k|, |p_j - p_k|$

$\hat{\delta}$ functions have width $\sim \frac{1}{\sigma_{s,D}} < \frac{1}{A} \approx \text{keV}$

↳ terms that vary slowly over such a momentum interval can be pulled out of the integral

$$\Rightarrow \mathcal{A} \propto \left(\bar{u}_B \gamma^\mu \frac{\not{p}_0 - \not{v}}{2p_0} \not{\epsilon} \mathcal{M}_\nu^S \mathcal{M}_\nu^D + O\left(\frac{m_j^2}{p_0^2}\right) \right)$$

$\sim p_0 + p_w$

$$\cdot \sum_j \mathcal{M}_{\alpha_j}^* \mathcal{M}_{\beta_j} \int \frac{d^3 p_j}{(2\pi)^3} e^{-i(E_{jT} - \vec{p}_j \cdot \vec{L})}$$

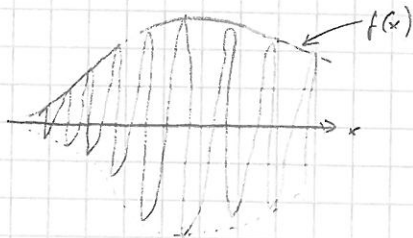
$$\cdot \hat{\delta}(p_\nu - p_k - p_w) \hat{\delta}(p_\nu - p_k - p_w)$$

Method of stationary phase:

$$\int dx f(x) e^{i\phi(x)}$$

function that varies significantly over interval Δx

can be large only if $f(x)$ has $O(1)$ variations over intervals $\approx \Delta x$
(otherwise, just oscillations average integral to zero)



Here, this means only region with $\vec{p}_\nu \parallel \vec{L}$ relevant \Rightarrow reduce int. to 1D

$$\text{Moreover, enforces } T \approx \frac{L}{E_{j0}} + O(\sigma_{s,D}) = \frac{L}{\sqrt{p_0^2 + m_j^2}}$$

Let us choose $T \equiv |L|$

$$\Rightarrow \mathcal{A} \propto \sum_j \mathcal{M}_{\alpha_j}^* \mathcal{M}_{\beta_j} \int dp_\nu e^{-i \frac{m_j^2 L}{2p_\nu}}$$

$$\cdot \hat{\delta}(p_\nu - p_k - p_w) \hat{\delta}(p_\nu - p_k - p_w)$$

$$\approx \sum_j \mathcal{M}_{\alpha_j}^* \mathcal{M}_{\beta_j} e^{-i \frac{m_j^2 L}{2p_0}}$$

$$E_{0j} = \sqrt{p_\nu^2 + m_j^2} \approx p_\nu + \frac{m_j^2}{2p_\nu}$$

$$\frac{m_j^2 L}{2p_\nu} = \frac{m_j^2 L}{2(p_0 + \sigma_{s,D})} \approx \frac{m_j^2 L}{2p_0} \frac{\sigma_{s,D}}{p_0}$$

$$\approx \frac{(1\text{eV})^2 \cdot \text{km}}{2 \cdot \text{MeV}} \frac{\text{keV}}{\text{MeV}} \ll 1$$

Probability:

$$P_{\alpha \rightarrow \beta} = \frac{1}{\mathcal{N}} |\mathcal{A}|^2$$

$$\propto \frac{1}{\mathcal{N}} \sum_{j,k} U_{\alpha j}^* U_{\beta k} U_{\alpha j} U_{\beta k}^* \int d\tilde{p}_j d\tilde{p}_k \frac{d\tilde{p}_j d\tilde{p}_k}{dE_j dE_k} e^{-i(E_j T - p_j L) + i(\tilde{E}_j T - \tilde{p}_j L)} \cdot \delta(p_\alpha - p_j - p_i) \delta(p_\alpha - p_k - p_i) \cdot \delta(\tilde{p}_\alpha - \tilde{p}_j - \tilde{p}_i) \delta(\tilde{p}_\alpha - \tilde{p}_k - \tilde{p}_i)$$

Assume continuous neutrino flux ("stationary source"), and assume we're not interested in when each ν is detected

$$\hookrightarrow \overline{P_{\alpha \beta}} = \frac{1}{\mathcal{N}} \int dT |\mathcal{A}|^2$$

$$= \frac{1}{\mathcal{N}} \sum_{j,k} U_{\alpha j}^* U_{\beta j} U_{\alpha k} U_{\beta k}^* \int dE_j e^{i(\sqrt{E_j^2 - m_j^2} - \sqrt{E_j^2 - m_i^2})L} \cdot \delta(\sqrt{E_j^2 - m_j^2} - p_j - p_i) \delta(\sqrt{E_j^2 - m_i^2} - p_\alpha - p_i) \cdot \delta(\sqrt{E_k^2 - m_k^2} - p_k - p_i) \delta(\sqrt{E_k^2 - m_i^2} - p_\alpha - p_i)$$

Use $(\sqrt{E_j^2 - m_j^2} - \sqrt{E_j^2 - m_i^2})L \approx -\frac{m_j^2 - m_i^2}{2E_j} L$

varies slowly over interval $[E_j, E_j + \epsilon_{j,p}]$

\rightarrow pull out of integral as const.

$$\Rightarrow \overline{P_{\alpha \beta}} = \sum_{j,k} U_{\alpha j}^* U_{\beta j} U_{\alpha k} U_{\beta k}^* \cdot \exp\left[-i \frac{\Delta m_{jk}^2}{2E_j} L\right]$$

$$P_{\alpha \rightarrow \beta} = |\mathcal{A}|^2 = \sum_{j,k} U_{\alpha j}^* U_{\beta k} U_{\alpha j} U_{\beta k}^* e^{-i(E_j - E_k)T + i(p_j - p_k)L}$$

Assume continuous neutrino flux ("stationary source"), and assume we're not interested in when each neutrino is detected

$$\hookrightarrow \overline{P_{\alpha \rightarrow \beta}} = \frac{1}{\mathcal{N}} \int dT |\mathcal{A}|^2$$

\int normalization

$$= \frac{1}{\mathcal{N}} \sum_{j,k} U_{\alpha j}^* U_{\beta k} U_{\alpha j} U_{\beta k}^* \cdot 2\pi \delta(E_j - E_k)$$

$$\cdot \exp\left[i(\sqrt{E_j^2 - m_j^2} - \sqrt{E_j^2 - m_i^2})L\right]$$

$$= \sum_{j,k} U_{\alpha j}^* U_{\beta k} U_{\alpha j} U_{\beta k}^*$$

$$\cdot \exp\left[-i \frac{\Delta m_{jk}^2}{2E} L\right]$$

Check normalization: $\sum_{\beta} \overline{P_{\alpha \beta}} = \sum_{j,k} U_{\alpha j}^* U_{\alpha k} \delta_{jk} e^{-i \frac{\Delta m_{jk}^2 L}{2E_j}}$

$$= 1$$

Consider 2-flavor case:

$$U = \begin{pmatrix} \cos \theta & \sin \theta \\ -\sin \theta & \cos \theta \end{pmatrix}$$

$$\begin{aligned} \Rightarrow P_{e \rightarrow \mu} &= |U_{e1}|^2 |U_{\mu 1}|^2 + |U_{e2}|^2 |U_{\mu 2}|^2 + U_{e1} U_{\mu 1} U_{e2} U_{\mu 2} \left(e^{-i \frac{\Delta m^2 L}{2E}} + e^{+i \frac{\Delta m^2 L}{2E}} \right) \\ &= \cos^2 \theta \sin^2 \theta \cdot 2 - \cos^2 \theta \sin^2 \theta \cdot 2 \cos \frac{\Delta m^2 L}{2E} \\ &= 2 \cdot \frac{1}{4} \sin^2 2\theta - 2 \cdot \frac{1}{4} \sin^2 2\theta \cos \frac{\Delta m^2 L}{2E} \\ &= \sin^2 2\theta \sin^2 \frac{\Delta m^2 L}{4E} \end{aligned}$$

2.3 3-flavor neutrino oscillations

3-flavor mixing matrix

$$\begin{pmatrix} \nu_e \\ \nu_\mu \\ \nu_\tau \end{pmatrix} = U \begin{pmatrix} \nu_1 \\ \nu_2 \\ \nu_3 \end{pmatrix}$$

Parameters:	General 3×3 matrix:	9 real params	9 phases
	Unitary: $\sum_j U_{\alpha j} ^2 = 1$	-3	
	$\sum_j U_{\alpha j} U_{\beta j}^* \stackrel{\alpha \neq \beta}{=} 0$	-3	-3
	Field redefinitions $\nu_j \rightarrow e^{i\phi_j} \nu_j$ $\nu_\alpha \rightarrow e^{i\phi_\alpha} \nu_\alpha$		-5 (a common rescaling of all ν_i, ν_j does not alter U)

\Rightarrow 3 real parameters ("mixing angles"), 1 phase

Convenient parameterization: $\theta_{12}, \theta_{13}, \theta_{23}, \delta$

$$\begin{aligned} c_{ij} &\equiv \cos \theta_{ij} \\ s_{ij} &\equiv \sin \theta_{ij} \end{aligned}$$

$$U = \begin{pmatrix} 1 & & & & & \\ & c_{23} & s_{23} & & & \\ & -s_{23} & c_{23} & & & \\ & & & 1 & & \\ & & & & c_{13} & s_{13} e^{-i\delta} \\ & & & & -s_{13} e^{i\delta} & c_{13} \end{pmatrix} \begin{pmatrix} c_{12} & s_{12} \\ -s_{12} & c_{12} \\ & & 1 \end{pmatrix}$$

In addition, $P_{\alpha \rightarrow \beta}$ depends on two independent "mass squared differences"

$$\Delta m_{21}^2 \equiv m_2^2 - m_1^2 \quad \text{and} \quad \Delta m_{31}^2 \equiv m_3^2 - m_1^2$$

2.4 Neutrino oscillation experiments

Neutrino sources

↳ see presentation

Neutrino detectors

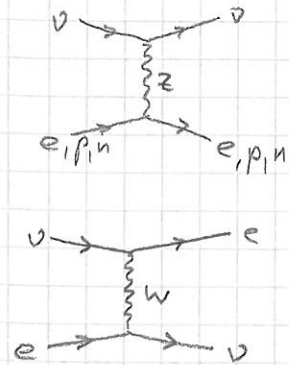
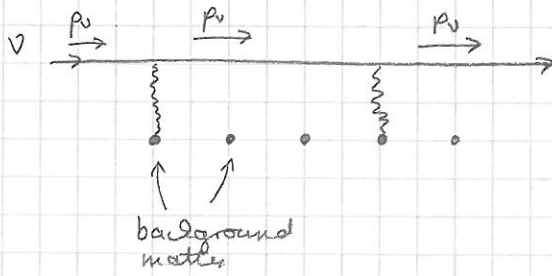
↳ see presentation

Neutrino oscillation experiments

↳ see presentation

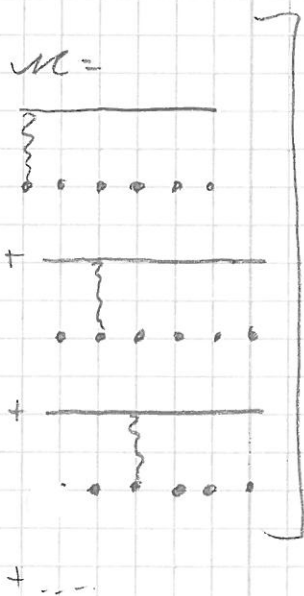
2.5 Neutrino oscillations in matter

Coherent forward scattering:



Exchange of force carrier (W, Z) with background matter, leaving energies and momenta unchanged.

Analogy: Photon travelling through glass



Crucial observation:

Interactions with different particles add up coherently

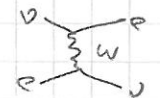
$$\hookrightarrow |\mathcal{M}|^2 \propto N^2 \cdot G_F^2$$

↑
number of matter particles

(cf. incoherent scattering: $|\mathcal{M}|^2 \propto N \cdot G_F^2$)

\Rightarrow enhancement by N

Coherent forward scattering creates a potential:

Consider 

$$\mathcal{M}_{\text{cf}} \equiv \frac{G_F}{\sqrt{2}} [\bar{e} \gamma^\mu (1 - \gamma^5) \nu_e] [\bar{\nu}_e \gamma^\mu (1 - \gamma^5)]$$

$$\stackrel{\text{Fierz transformation}}{\Rightarrow} \frac{G_F}{\sqrt{2}} [\bar{e} \gamma^\mu (1 - \gamma^5) e] [\bar{\nu}_e \gamma^\mu (1 - \gamma^5)]$$

Treat electrons as fixed background field
 \Rightarrow take expectation values $\langle \cdot \rangle_e$

$$\langle \mathcal{L}_{\text{eff}} \rangle_e = \frac{1}{i^4} \langle \bar{e} \gamma^\mu (1 - \gamma^5) e \rangle_e \cdot [\bar{v}_e \gamma^\mu (1 - \gamma^5) v_e]$$

Remember:

$$1_4 = \begin{pmatrix} 1 & 0 & 0 & 0 \\ 0 & 1 & 0 & 0 \\ 0 & 0 & 1 & 0 \\ 0 & 0 & 0 & 1 \end{pmatrix}$$

tr $\sigma^{\mu\nu} = 2g^{\mu\nu}$

- $\langle \bar{e} \gamma^0 e \rangle \sim \langle e^\dagger e \rangle = n_e$
 \uparrow e number density
- $\langle \bar{e} \vec{\gamma} e \rangle \sim \langle \vec{v}_e \rangle = 0$ in matter at rest
- $\langle \bar{e} \gamma^0 \gamma^5 e \rangle \sim \langle \frac{\vec{\sigma} \cdot \vec{p}_e}{E} \rangle = \text{avg. helicity} = 0$
 for unpolarized matter / matter at rest
- $\langle \bar{e} \vec{\gamma} \gamma^5 e \rangle \sim \langle \vec{\sigma} \rangle = \text{average spin} = 0$
 for unpolarized matter

$$\text{Use } e = \int \frac{d^3 p}{(2\pi)^3} \frac{1}{\sqrt{2E_p}} \sum_{\text{spins } s} \left(a_{\vec{p}}^s u^s(p) e^{-ipx} + b_{\vec{p}}^{s\dagger} v^s(p) e^{ipx} \right)$$

$\uparrow = \sqrt{p^2 + m_e^2}$ \uparrow particle annihilation operator \uparrow anti-particles creation operator

$$|e^-(\vec{p}, s)\rangle = \sqrt{2E_p} a_{\vec{p}}^{s\dagger} |0\rangle$$

Assume $\langle \bar{e} e \rangle$ follows p -distribution $f(\vec{p})$; $\int \frac{d^3 p}{2E_p} f(\vec{p}) = n_e$
 \uparrow number density

$$\Rightarrow \langle \bar{e} \gamma^\mu (1 - \gamma^5) e \rangle_e = \frac{1}{i^4} \sum_s \int \frac{d^3 p}{2E_p} f(\vec{p}) \langle e^-(\vec{p}, s) | \bar{e} \gamma^\mu (1 - \gamma^5) e | e^-(\vec{p}, s) \rangle$$

$$\left\{ a_{\vec{p}}^s, a_{\vec{p}'}^{s'\dagger} \right\} = (2\pi)^3 \delta^{(3)}(\vec{p} - \vec{p}') \delta_{ss'}$$

$$\stackrel{\text{spin average}}{\uparrow} = \frac{1}{2} \sum_s \int \frac{d^3 p}{2E_p} f(\vec{p}) \cdot \bar{u}^s(p) \gamma^\mu (1 - \gamma^5) u^s(p)$$

$$= \frac{1}{2} \int \frac{d^3 p}{2E_p} f(\vec{p}) \text{tr}((\not{p} + m_e) \gamma^\mu (1 - \gamma^5))$$

$$= \frac{1}{2} \int \frac{d^3 p}{2E_p} f(\vec{p}) \frac{\text{tr} \not{p} \gamma^\mu}{= 4p^\mu}$$

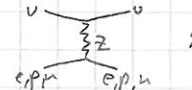
$$= \begin{cases} n_e & \text{for } \mu = 0 \\ 0 & \text{for } \mu = 1, 2, 3 \end{cases} \quad \text{assuming } f(\vec{p}) = f(|\vec{p}|)$$

$$\left. \begin{aligned} \text{tr } \gamma^\mu &= 0 \\ \text{tr } \gamma^\mu \gamma^5 &= 0 \\ \text{tr } \gamma^\mu \gamma^\nu \gamma^5 &= 0 \end{aligned} \right\}$$

$$\Rightarrow \langle \mathcal{L}_{\text{eff}} \rangle = \frac{G_F}{\sqrt{2}} n_e \bar{\nu}_e \gamma^0 (1 - \gamma^5) \nu_e$$

q. similar structure:
em potential A_μ
 $\hookrightarrow \mathcal{L} \supset \bar{e} \gamma^\mu e A_\mu$

$$= \frac{\sqrt{2} G_F n_e}{\equiv V_{CC}} \bar{\nu}_{eL} \gamma^0 \nu_{eL}$$

Similarly, potential from  ;

$$V_{NC} = -\frac{1}{2} \cdot \sqrt{2} G_F n_e$$

affects all neutrino flavors

Note: Potentials flip sign under CP (deeper reason: $\bar{\nu}_{eL} \gamma^0 \nu_{eL} = -(\bar{\nu}_{eL})^c \gamma^0 (\nu_{eL})$)

In flavor basis $\Psi \equiv \begin{pmatrix} a |\nu_e\rangle \\ b |\nu_\mu\rangle \end{pmatrix}$, $a^2 + b^2 = 1$

$$P = \left| \Psi^\dagger U \begin{pmatrix} e^{i\hat{p}_\mu L} \\ e^{i\hat{p}_\nu L} \end{pmatrix} U^\dagger \Psi \right|^2$$

$$\begin{aligned} &= \left| \Psi^\dagger e^{i\hat{p}L} \Psi \right|^2 \\ &= \left| \Psi^\dagger e^{i \underbrace{\hat{H}^2 - \hat{p}^2}_{\hat{H}_{\text{kin}}}} \Psi \right|^2 \\ &\quad \hat{H}_{\text{kin}} = \hat{H} - \hat{V} \\ &\approx \left| \Psi^\dagger \exp \left[i \left(\hat{H} + \frac{\hat{H}^2}{2\hat{H}} - \hat{V} \right) \right] \Psi \right|^2 \end{aligned}$$

\uparrow in flavor basis: $\begin{pmatrix} V_{CC} - V_{NC} & \\ & -V_{NC} \end{pmatrix}$

Consider phase $\phi = \hat{p} \cdot L$

$$= \sqrt{(\hat{H} - \hat{V})^2 - \hat{H}^2}$$

Expand small \downarrow

$$\approx \hat{H} - \frac{\hat{H}^2}{2\hat{H}} - \hat{V}$$

Need to determine states of definite momentum

$$\hookrightarrow \text{Diagonalize } \hat{H} - \frac{\hat{H}^2}{2\hat{H}} - \hat{V} = E \cdot 1 - U \begin{pmatrix} \frac{m_1^2}{2E} & \\ & \frac{m_2^2}{2E} \end{pmatrix} U^\dagger - \begin{pmatrix} V_{CC} & \\ & V_{NC} \end{pmatrix}$$

$$= \left(E - \frac{m_1^2}{4E} - \frac{m_2^2}{4E} - V_{NC} \right) \cdot \begin{pmatrix} \cos \theta & \sin \theta \\ -\sin \theta & \cos \theta \end{pmatrix} \begin{pmatrix} \frac{\Delta m^2}{4E} \\ \frac{\Delta m^2}{4E} \end{pmatrix} \begin{pmatrix} \cos \theta & -\sin \theta \\ \sin \theta & \cos \theta \end{pmatrix} - \begin{pmatrix} V_{CC} & \\ & 0 \end{pmatrix}$$

$\begin{matrix} \equiv s & \equiv c \end{matrix}$

see exercise, mention only outline of calculation and result

$$\begin{aligned}
 &= \left(E - \frac{m_1^2}{4E} - \frac{m_2^2}{4E} - V_{Nc} \right) - \begin{pmatrix} c & s \\ -s & c \end{pmatrix} \begin{pmatrix} -c & s \\ s & c \end{pmatrix} \frac{\Delta m^2}{4E} - \begin{pmatrix} V_{cc} \\ 0 \end{pmatrix} \\
 &= \quad \quad \quad - \begin{pmatrix} -c^2 + s^2 & 2sc \\ 2sc & -s^2 + c^2 \end{pmatrix} \frac{\Delta m^2}{4E} - \quad \quad \quad \\
 &= \quad \quad \quad - \begin{pmatrix} \frac{\Delta \cos 2\theta}{2} + \frac{\Delta \sin 2\theta}{2} \\ \frac{\Delta \sin 2\theta}{2} & \frac{\Delta \cos 2\theta}{2} \end{pmatrix}
 \end{aligned}$$

$$\Rightarrow \left(V_{cc} - \frac{\Delta \cos 2\theta}{2} - \lambda \right) \left(\frac{\Delta \cos 2\theta}{2} - \lambda \right) - \frac{\Delta^2 \sin^2 2\theta}{4} = 0$$

$$\Leftrightarrow \lambda^2 - \lambda \cdot V_{cc} - \frac{\Delta^2 \cos^2 2\theta}{4} + V_{cc} \frac{\Delta \cos 2\theta}{2} - \frac{\Delta^2 \sin^2 2\theta}{4} = 0$$

$$\begin{aligned}
 \lambda &= \frac{V_{cc}}{2} \pm \sqrt{V_{cc}^2 + \Delta^2 - 2V_{cc} \Delta \cos 2\theta} \\
 &= \frac{V_{cc}}{2} \pm \sqrt{(V_{cc} - \Delta \cos 2\theta)^2 + \Delta^2 \sin^2 2\theta}
 \end{aligned}$$

Eigenvectors: $\begin{pmatrix} \text{I} \\ \text{II} \end{pmatrix} \begin{pmatrix} -\frac{\Delta}{2} \cos 2\theta + V_{cc} & \frac{\Delta}{2} \sin 2\theta \\ \frac{\Delta}{2} \sin 2\theta & \frac{\Delta}{2} \cos 2\theta \end{pmatrix} \begin{pmatrix} a \\ b \end{pmatrix} = \frac{1}{2} V_{cc} \begin{pmatrix} a \\ b \end{pmatrix} \pm \sqrt{\dots} \begin{pmatrix} a \\ b \end{pmatrix}$

$$\cos 2\theta \cdot \text{I} - \sin 2\theta \cdot \text{II} : -\frac{\Delta}{2} a + V_{cc} \cos 2\theta a = \frac{1}{2} (V_{cc} \pm Q) (\cos 2\theta a - \sin 2\theta b)$$

$$\Rightarrow \sin 2\theta \cdot b \cdot \frac{1}{2} (V_{cc} \pm Q)$$

$$= \left[\cos 2\theta \cdot \frac{1}{2} (V_{cc} \pm Q) + \frac{\Delta}{2} - V_{cc} \cos 2\theta \right] a$$

$$= \left[\frac{\Delta}{2} + \frac{1}{2} \cos 2\theta (-V_{cc} \pm Q) \right] a$$

$$\Rightarrow \tan \theta_{\text{eff}} \equiv \left. \frac{-b}{a} \right|_{\text{II}} = - \frac{\frac{\Delta}{2} + \frac{1}{2} \cos 2\theta (-V_{cc} - Q)}{\sin 2\theta \cdot \frac{1}{2} (V_{cc} - Q)}$$

$$\sin 2\theta_{\text{eff}} = \frac{2 \tan \theta_{\text{eff}}}{1 + \tan^2 \theta_{\text{eff}}}$$

$$= -2 \frac{\frac{\Delta}{2} + \cos 2\theta (-V_{cc} - Q)}{\sin 2\theta (V_{cc} - Q) \left[1 + \left(\frac{\frac{\Delta}{2} + \cos 2\theta (-V_{cc} - Q)}{\sin 2\theta (V_{cc} - Q)} \right)^2 \right]}$$

$$\begin{aligned}
&= -2 \frac{(\Delta - \cos 2\theta (V_{cc} + G)) \sin 2\theta (V_{cc} - G)}{\sin^2 2\theta (V_{cc} - G)^2 + (\Delta - \cos 2\theta (V_{cc} + G))^2} \\
&= -2 \frac{\sin 2\theta [\Delta V_{cc} - \Delta G - \cos 2\theta (V_{cc}^2 - G^2)]}{\sin^2 2\theta (V_{cc}^2 - 2V_{cc}G + G^2) + \Delta^2 + \cos^2 2\theta (V_{cc}^2 + G^2 + 2V_{cc}G) - 2\Delta \cos 2\theta (V_{cc} + G)} \\
&= -2 \sin 2\theta \frac{\Delta [V_{cc} - G - \cos 2\theta (-\Delta + 2V_{cc} \cos 2\theta)]}{V_{cc}^2 + G^2 + 2V_{cc}G (\cos^2 2\theta - \sin^2 2\theta) + \Delta^2 - 2\Delta \cos 2\theta (V_{cc} + G)} \\
&= -2 \sin 2\theta \cdot \Delta \frac{V_{cc} (\sin^2 2\theta - \cos^2 2\theta) - G + \Delta \cos 2\theta}{2G^2 + 2V_{cc}G (\cos^2 2\theta - \sin^2 2\theta) - 2G \Delta \cos 2\theta} \\
&= \frac{\Delta \cdot \sin 2\theta}{G}
\end{aligned}$$

Result: $\frac{\Delta m^2_{\text{eff}}}{2E} = \sqrt{(\sqrt{2} G_F n_e - \frac{\Delta m^2}{2E} \cos 2\theta)^2 + (\frac{\Delta m^2}{2E})^2 \sin^2 2\theta}$

$$\sin^2 2\theta_{\text{eff}} = \frac{\sin 2\theta \cdot \frac{\Delta m^2}{2E}}{\sqrt{(\sqrt{2} G_F n_e - \frac{\Delta m^2}{2E} \cos 2\theta)^2 + (\frac{\Delta m^2}{2E})^2 \sin^2 2\theta}}$$

At the MSW (Mikheyev-Smirnov-Wolfenstein) resonance,

$$\frac{\Delta m^2}{2E} \cos 2\theta = \sqrt{2} G_F n_e$$

the effective mixing θ_{eff} is maximal.

3. Neutrino anomalies

↳ see presentation